Glossary 2010
ablation
achromatic lens
albedo
alignment
altitude
appulse
aphelion
arcminute
arcsecond
ascending node
asteroid
astronomical unit
asynchronous
axis
azimuth
bandpass filter
binary star
candela
Cassegrain telescope
cataclysmic variable
celestial equator
celestial poles
Charge
chromatic aberration
conjunction
contrast
constellation
black hole a region of space-time that cannot be seen by distant observers because light is trapped by a strong gravitational field
erosion of an object (generally a meteorite) by the friction generated when it passes through the Earth's atmosphere
a compound lens whose elements differ in refractive constant in order to minimize chromatic aberration
the ratio of the amount of light reflected from a surface to the amount of incident light
the adjustment of an object in relation with other objects
the angular distance of a celestial body above or below the horizon
a penumbral eclipse of the Moon
the point on its orbit where the Earth is farthest from the Sun
one sixtieth of a degree of angular measure
one sixtieth of an arcminute, or $1 / 3600$ of a degree
in the orbit of a Solar System body, the point where the body crosses the ecliptic from south to north
a small rocky body that orbits a star — in the Solar System, most asteroids lie between the orbits of Mars and Jupiter
mean distance between the Earth and the Sun
in connection with orbital mechanics, refers to objects that pass overhead at different times of the day; does not move at the same speed as Earth's rotation
theoretical straight line through a celestial body, around which it rotates
the direction of a celestial body from the observer, usually measured in degrees from north
a device for suppressing unwanted frequencies without appreciably affecting the desired frequencies
two stars forming a physically bound pair under their mutual gravitational attraction - the stars move in elliptical orbits about their common centre of mass the SI base unit of luminous intensity
a telescope devised by Cassegrain in which an auxiliary convex mirror reflects the magnified image, upside down, through a hole in the centre of the main objective mirror
a star in which the brightness increases suddenly because of an explosive event
projection of the Earth's equator as a line across the sky (for an observer on the equator, such a line would pass through the zenith)
the two points at which the Earth's axis of rotation, if extended, would intersect the celestial sphere
the fundamental property of a particle that causes it to be affected by the electromagnetic force
introduction of spurious colors by a lens, attenuated by the introduction of corrective elements into a compound lens
the phenomenon in which two bodies have the same apparent celestial longitude or right ascension as viewed from a third body
the difference in visual properties that makes an object distinguishable from other objects and the background
a group of celestial bodies (usually stars) that appear to form a pattern in the

| coordinates | quantities that provide references for locations in space and time |
| :---: | :---: |
| corona | outermost atmosphere of the Sun |
| cosmic rays | high-speed particles that reach the Earth from outside the Solar System |
| culminate | to reach the highest point above an observer's horizon |
| declinati | angular distance above or below the celestial equator - one of the coordinates, with right ascension, that defines the position of a heavenly |
|  |  |
| deferent | in the Ptolemaic system, the planets are assumed to move in a small circle, called an epicycle, which in turn moves along a larger circle called a |
|  | deferent. |
| descending node | in the orbit of a Solar System body, the point where the body crosses the ecliptic from north to south |
| ecliptic | the apparent path that the Sun traces out in the sky during the year, so named because eclipses occur when the full or new Moon is very close to this path |
|  | of the Sun |
| elongation | a planet's elongation is the angle between the Sun and the planet, as viewed from Earth |
| ephemeris | a table of values that gives the positions of astronomical objects in the sky at a given time or times; plural = ephemerides |
| epicycle | in the Ptolemaic system, the planets are assumed to move in a small circle, called an epicycle (see deferent) |
| exit pupil | a virtual aperture in an optical system; an image of the objective element(s) as produced by a binocular or telescope eyepiece |
| flux | in magnetism, the total number of lines of magnetic force passing through a specified loop; a measure of the amount of power or radiation received per unit time per unit area; |
| galaxy | vast system of celestial objects, typically consisting of between $10^{6}$ and $10^{12}$ stars, plus interstellar gas and dust |
| gegenshein | a faint oval patch of light visible from Earth only at certain times of the year, opposite the Sun |
| geocentri | with reference to, or pertaining to, the centre of the Earth |
| geodesic | a path or line of shortest distance joining two points in space (or space-time) |
| gravity (gravitation) | the universal ability of all material objects to attract each other; its force is directly proportional to the mass of each object, and decreases by the square of the distance separating the objects involved |
| Gregorian calendar | the calendar introduced by Pope Gregory XIII in 1582 to replace the Julian calendar; the calendar now used as the civil calendar in most countries. |
| heiligenschien | an optical phenomenon that creates a bright spot around the shadow of the viewer's head |
| heliocentric | a cosmological system in which the Sun is at (or near) the central point |
| Hertzsprung-Russell |  |
| Diagram | a plot of stellar color, temperature, or spectral type versus stellar luminosity |
| illuminance | the total luminous flux incident on a surface, per unit area, i.e. one lux is an illumination of one lumen per square metre |
| immersion | the disappearance of a celestial body due to eclipse or occultation |
| inclination | the angle between one plane and another; the (equatorial) inclination of a planet is the angle between the plane of its equator and that of its orbit; the inclination of the orbit of a planet in the Solar System other than Earth is the angle between the plane of that orbit and the ecliptic |
| Lagrangian points | five points in the orbital plane of two massive objects in orbits aro |

$\left.\begin{array}{ll} & \begin{array}{l}\text { common centre of gravity, where a third body of negligible mass can remain } \\ \text { in equilibrium } \\ \text { angular distance on the celestial sphere measured north or south of the } \\ \text { ecliptic along the great circle passing through the poles of the ecliptic and the }\end{array} \\ \text { celestial object } \\ \text { distance travelled at the speed of light after one Earth-year: } 9.46 \text { million } \\ \text { million km } \\ \text { angular distance measured along the Earth's equator from the } \\ \text { Greenwich meridian to the meridian of a geographic location }\end{array}\right\}$
$\left.\begin{array}{ll}\text { radiant } & \begin{array}{l}\text { the point in the sky from which a meteor shower appears to emanate } \\ \text { redshift } \\ \text { the shift of spectral lines toward longer wavelengths in the spectrum of } \\ \text { a receding source of radiation } \\ \text { the change in direction of travel (bending) of a light ray as it passes } \\ \text { obliquely through the atmosphere; in the case of a lens, any ray as it } \\ \text { passes from one medium into another of greater or lesser density } \\ \text { the theory of how motion and gravity affect the properties of time and } \\ \text { space } \\ \text { refraction }\end{array} \\ \text { relativity } & \text { when two orbiting bodies exert a regular, periodic gravitational } \\ \text { influence on each other } \\ \text { angular distance on the celestial sphere measured eastward along the } \\ \text { celestial equator from the equinox to the hour circle passing through }\end{array}\right\}$

